



On Mathematical Analysis LRS Bianchi Type II String Cosmological Models with Viscosity Distribution in Modified General Relativity

*R. K. Dubey¹, Shishir Kumar Srivastava² & J .N. Mishra³

¹Department of Mathematics Govt. Model Science College, Rewa (M.P.) India

²Department of Mathematics Ganpat Sahai P.G. College, Sultanpur (U.P.) India

³Department of Mathematics Ganpat Sahai P.G. College, Sultanpur (U.P.) India

Email: shishir9918825052@gmail.com

Abstract

We consider locally rotationally symmetric Bianchi type II string cosmological model in the presence of viscous fluid. To solve the Einstein's field equations for LRS Bianchi type II space time has been obtained under the assumption $\rho = K_{m_1}\eta$, here ρ is the energy density, η is the string tension density and K_{m_1} is a constant. We have also used a condition that, the scalar expansion is proportional to the shear, to get determinate solution in terms of cosmic time t . Some physical and model geometric behavior of the models is discussed.

Keywords: LRS Bianchi type II, massive string, viscous fluid.

MSC2010 CLASSIFICATIONS: 83C05, 83C15.

PACS Number: 98.80cq, 04.20.-q

References

- 1- Bali, R. and Dave, S. (2001). Bianchi type IX string cosmological model in general relativity Pramana. J. Phys. 56, 53-518.
- 2- Bali, R. and Anjali (2003). Bianchi type IX magnetized string cosmological model in general relativity Astrophys. space sci 288, 399-405.
- 3- Bali, R. and singh, D.K. (2005) Bianchi type V bulk viscous fluid dust cosmological model in general relativity Astophys. spce. Sci. 300, 387-394.
- 4- Bali, R. Pareek, U. K. and Pradhan, A. (2007). Bianchi type I massive string magnetized barotropic perfect fluid cosmological model in general relativity, chin phys. Lett 24; 2455-2458.
- 5- Bali, R. and Upadhyaya, R. D. (2003). LRS Bianchi type I string dust magnetized cosmological models Astrophys. spce Sci 283; 97-108.
- 6- Bali, R., and Anjali (2006). Bianchi type I magnetized string cosmological model in general relativity Astrophys. space Sci. 302, 201-205.
- 7- Banerjee, A., Dutta, S. B. and Sanyal, A. K. (1986). Bianchi type III Cosmologicla model with viscous fluid Gen. Relativity and Gravitation 18, 461-477.
- 8- Kibble, T. W. B. (1980). Some implications of a cosmological phase- transition, phys. Ref. 67, 183- 1999.
- 9- Kibble, T. W. B. (1976). Topology of cosmic domains and strings, J. Phys. A: math Gen 9, 1387-1398.
- 10- Letelier P. S. (1983). strings cosmologeis, Phys. Rev. D28, 2414-2419.
- 11- Pradhan, A., Amirhashchi, H. and Mahanta, K. L. (2007). Five dimentional LRS Bianchi type I string cosmological model in saez and Ballester theory Astrophys. Space Sci 312, 321-424.
- 12- Stachel, J. (1980). Thickening the string I the string perfect dust Phys. Rev. D21, 2171-2181.
- 13- Singh, T. and Agrawal, A. K. (1997). Bianchi type II, VIII, IX in certain new theories of gravitation Astrophys. space sc. 191, 61- 88.
- 14- Tyagi, A. and Sharma, K. (2010). Bianchi type II Bulk viscous string cosmological models in general relativity, International Journal of theoritical Phys 49(8), 1712-1718.
- 15- Vilenkin, A. (1981). Cosmic strings phys. Rev D.24, 2082-2089.
- 16- Wang, X. X. (2004). Bianchi type I, string cosmological models with bulk viscosity and magnetic field Astrophys. Space Sci. 293, 933-944.
- 17- Wang, X.X. (2005). Kantowski- Sachs string cosmological models with bulk viscosity in general relativity Astrophys. Space Sci 298, 433-440.
- 18- Wang, X. X. (2006). Bianchi type III string cosmological models with bulk viscosity and magnetic field chin, phys Lett. 23; 1702- 1704.
- 19- Yadav, M. K. Pradhan, A. and Rai, A. (2007). Some magnetized bulk viscous string cosmological models in general relativity, Astrophys. Space sci. 317; 423-429.
- 20- Yadav, M. K., Pradhan, A. and Rai, A. (2007). Some Bianchi type III string cosmological models with bulk viscosity, Int. J. Theor phys 46, 2677-2687.
- 21- Zel'dovich, ya, B. Kokzarev, I ya and Okun, L. B. (1975). Cosmological consequences of a spontaneous breakdown of a discrete symmetry, Zn. Sov. Phys. JETP 40, 1-5.